Note

Chebyshev Series Approximations for the Zeros of the Bessel Functions

1. INTRODUCTION

The zeros of the Bessel functions $J_{\nu}(x)$ have important applications in mathematical physics and applied mathematics. We denote the sth zero of $J_{\nu}(x)$ by $j_{\nu,s}$. Several approximations, asymptotic expansions or bounds for the zeros of the Bessel functions exist (see [1, 2, 5, 7, 13]. Especially McMahon's expansion for large zeros, (see Olver [7] or Abramowitz and Stegun [1]), Olver's asymptotic expansion for large orders, and Olver's uniform asymptotic expansions (see Olver [7]) are interesting formulas. Unfortunately, they are not applicable when s and ν are small. For that case, which is of frequent occurrence, we have then to use iterative methods for the computation of $j_{\nu,s}$. An excellent computer program, based on an iterative Newton process, is given by Temme [11].

The iterative computation of $j_{v,s}$ is very time-consuming, because it requires the evaluation of $J_v(x)$. For some applications, we need more efficient formulae. For example, for the estimation of the sth zero, $\xi_s^{(\alpha,n)}$ of the generalized Laguerre polynomial $L_n^{(\alpha)}(x)$, Tricomi [12] has given the following formula

$$\xi_{s}^{(\alpha,n)} = \frac{j_{\alpha,s}^{2}}{4k_{n}} \left[1 + \frac{2(\alpha^{2}-1) + j_{\alpha,s}^{2}}{48k_{n}^{2}} \right] + O(n^{-5}), \tag{1}$$

where $k_n = n + (\alpha + 1)/2$.

A similar formula exists for estimating the zeros of Gegenbauer polynomials [12]. Zeros of these orthogonal polynomials are the abscissae of Gaussian quadrature formulas (Stroud and Secrest [10]). The efficiency of numerical software for constructing Gaussian quadrature formulae is affected by the method for the computation of the abscissae. The most efficient methods are higher order iterative methods, especially when accurate starting values are easily available. Formula (1) gives very accurate approximations for the zeros of $L_n^{(\alpha)}(x)$, (especially for the smallest zeros). But the practical usefulness of (1) depends strongly on the existence of an efficient method for the computation of $j_{\alpha,s}$.

The purpose of this note is to present approximations for $j_{v,s}$ in the region of the small v- and s-values. By using these new approximations, or McMahon's asymptotic expansions, or Olver's uniform asymptotic expansions (depending on the values of v and s), we are able to calculate $j_{v,s}$ to at least 12 decimal figures, in the whole region $v \ge -1$, $s \ge 1$. Olver's uniform asymptotic expansions for $j_{v,s}$ are very powerful [7].

The first four terms yield twelve-figure accuracy when $v \ge 5$. McMahon's expansion, truncated after 8 terms yields also at least twelve-figure accuracy when s > 6 and v < 5 (the explicit expression for the coefficients of this expansion are given by Olver [7]). For $v \ge 5$, the range of applicability of these expansions partially overlap.

The gap left by Olver's and McMahon's expansions is partially closed by Chebyshev series expansions presented by Németh [6], which yield fifteen figure accuracy for $0 \le v \le \infty$ and $1 \le s \le 10$, and, by rational approximations given by Piessens [8], which are valid only for integer values of v.

Because for the important region -1 < v < 0, $1 \le s \le 6$, no accurate expansions or approximations for $j_{v,s}$ are known, we present here Chebyshev series approximations for $j_{v,s}$, s = 1, 2, 3, 4, 5, and 6 as functions of v on the interval [-1, 5].

TABLE I

Coefficients of the Approximation for $j_{r,1}$ (Formula (3))

k	$c_k^{(1)}$	k	
0	5.767950632456	0	
1	0.767665211539	1	
2	-0.086538804759	2	-
3	0.020433979038	3	
4	-0.006103761347	4	-
5	0.002046841322	5	
6	0.000734476579	6	-
7	0.000275336751	7	
8	-0.000106375704	8	-
9	0.000042003336	9	
10	-0.000016858623	10	-
11	0.000006852440	11	
12	-0.000002813300	12	-
13	0.000001164419	13	
14	0.000000485189	14	-
15	0.00000203309	15	
16	-0.00000085602	16	-
17	0.00000036192	17	
18	-0.00000015357	18	-
19	0.00000006537	19	
20	0.00000002791	20	-
21	0.00000001194	21	
22	-0.00000000512	22	-
23	0.00000000220	23	
24	-0.0000000095	24	_
25	0.00000000041	25	
26	-0.0000000018	26	
27	0.0000000008	27	
28	-0.00000000003	28	-
29	0.00000000001	29	

TABLE II

Coefficients of the Approximation for j_r , (Formula (4), s = 2)

k	$c_{k}^{(2)}$
0	16.526388664614
1	4.209200330779
2	-0.164644722483
3	0.039764618826
4	0.011799527177
5	0.003893555229
6	-0.001369989689
7	0.000503054700
8	-0.000190381770
9	0.000073681222
10	-0.000029010830
11	0.000011579131
12	-0.000004672877
13	0.000001903082
14	-0.000000781030
15	0.000000322648
16	-0.000000134047
17	0.00000055969
18	-0.00000023472
19	0.00000009882
20	-0.000000004175
21	0.00000001770
22	-0.000000000752
23	0.00000000321
24	-0.00000000137
25	0.00000000059
26	0.00000000025
27	0.00000000011
28	-0.000000000005
9	0.000000000002

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2. Chebyshev Series Approximations for $j_{\nu,s}$

In [9] it is shown that

$$j_{\nu,1} = (\nu+1)^{1/2} \left[2 + (\nu+1)/2 + O(\nu+1) \right], \quad \nu \to -1.$$
⁽²⁾

In order to take account of this asymptotic behaviour for $v \to -1$, $j_{v,1}$ is approximated by

$$j_{\nu,1} = (\nu+1)^{1/2} \sum_{k=0}^{N_1} c_k^{(1)} T_k \left(\frac{\nu-2}{3}\right), \qquad -1 \le \nu \le 5,$$
(3)

where the prime indicates that the first term is taken with factor $\frac{1}{2}$. For the following zeros, we use the approximation

$$j_{\nu,s} = \sum_{k=0}^{N_s} c_k^{(s)} T_k \left(\frac{\nu - 2}{3} \right), \qquad -1 \leqslant \nu \leqslant 5, \, s = 2, \, 3, \, 4, \, 5, \, 6.$$
(4)

TABLE III

TABLE IV

Coefficients of the Approximation for $j_{\nu,3}$ (Formula (4), s = 3)

Coefficients of the Approximation for $j_{r,4}$ (Formula (4), s = 4)

k	$c_{k}^{(3)}$	k	C ⁽⁴⁾
0	22.987742904346	0	29.378073011861
1	4.317988625384	1	4.387437455306
2	-0.130667664397	2	-0.109469595763
3	0.023009510531	3	0.015359574754
4	-0.004987164201	4	-0.002655024938
5	0.001204453026	5	0.000511852711
6	0.000310786051	6	-0.000105522473
7	0.000083834770	7	0.000022761626
8	-0.000023343325	8	-0.000005071979
9	0.000006655551	9	0.000001158094
10	0.000001932603	10	-0.00000269480
11	0.00000569367	11	0.00000063657
12	-0.000000169722	12	-0.00000015222
13	0.000000051084	13	0.00000003677
14	-0.00000015501	14	0.00000000896
15	0.00000004736	15	0.00000000220
16	-0.00000001456	16	-0.00000000054
17	0.00000000450	17	0.00000000013
18	-0.00000000140	18	-0.00000000003
19	0.00000000043	19	0.00000000001
20	0.00000000014		
21	0.00000000004		

$tor J_{r,5}$ (Formula (4), $s = 5$)		for $J_{r,6}$ (Formula (4), $s = 6$)		
k	$c_k^{(5)}$	k	$C_{k}^{(6)}$	
0	35.733765742756	0	42.069568616175	
1	4.435717974422	1	4.471319438161	
2	-0.094492317231	2	-0.083234240394	
3	0.011070071951	3	0.008388073020	
4	0.001598668225	4	-0.001042443435	
5	0.000257620149	5	0.000144611721	
6	-0.000044416219	6	-0.000021469973	
7	0.000008016197	7	0.000003337753	
8	-0.000001495224	8	-0.00000536428	
9	0.00000285903	9	0.00000088402	
10	-0.00000055734	10	-0.00000014856	
11	0.00000011033	11	0.00000002536	
12	-0.00000002212	12	-0.00000000438	
13	0.00000000448	13	0.00000000077	
14	-0.00000000092	14	-0.0000000014	
15	0.00000000019	15	0.00000000002	
16	-0.00000000004			

TABLE V

Coefficients of the Approximation for $j_{v,s}$ (Formula (4), s = 5)

The coefficients $c_k^{(s)}$ are computed numerically using an algorithm proposed by Gentleman [3]. The values of $j_{r,s}$ required for this algorithm, are calculated using an high-order iterative formula [4]. The values of N_s in (3) and (4) are chosen sufficiently large so that the error of the approximation is smaller than 10^{-12} . The values of the coefficients $c_k^{(s)}$ are presented in Tables 1–6.

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TABLE VI Coefficients of the Approximation

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